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Exploring Finite-Time Singularities and Onsager's Conjecture with Endpoint Regularity in the Periodic Navier Stokes Equations

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Abstract

It has recently been proposed by the author of the present work that the periodic NS equations (PNS) with high energy assumption can breakdown in finite time but with sufficient low energy scaling the equations may not exhibit finite time blowup. This article gives a general model using specific periodic special functions, that is degenerate elliptic Weierstrass P functions whose presence in the governing equations through the forcing terms simplify the PNS equations at the centers of cells of the 3-Torus. Satisfying a divergence free vector field and periodic boundary conditions respectively with a general spatio-temporal forcing term f which is smooth and spatially periodic, the existence of solutions which blowup in finite time for PNS can occur starting with the first derivative and higher with respect to time. P. Isett (2016) has shown that the conservation of energy fails for the 3D incompressible Euler flows with Hölder regularity below $1/3$. (Onsager's second conjecture) The endpoint regularity in Onsager's conjecture is addressed, and it is found that conservation of energy occurs when the Hölder regularity is exactly $1/3$. The endpoint regularity problem has important connections with turbulence theory. Finally very recent developed new governing equations of fluid mechanics are proposed to have no finite time singularities.

Index terms—

1 I. Introduction to the Periodic Navier Stokes Equations

The Navier-Stokes equations are useful because they describe the physics of many phenomena of scientific and engineering interest. They may be used to model the weather, ocean currents, pipe flows and heat exchangers and air flow around a wing. The Navier-Stokes equations, in their full and simplified forms, help with the design of aircraft and automobiles, hemodynamics, the design of power stations, the analysis of pollution and fuel emissions and many other things.

In 1845, Stokes had derived the equation of motion of a viscous flow by adding Newtonian viscous terms and finalized the Navier-Stokes equations, which have now been used for almost two centuries. There are only a few studies to find how to understand the physical meaning of the viscous terms in NS equations. As is well known, Stokes had three assumptions: 1. The force on fluids is the stationary pressure when the flow is stationary. 2. Fluid viscosity is isotropic. 3. Fluid flow follows Newton's law that fluid stress and strain have linear relations. These assumptions lead to the NSE. In [1], since the regular NS equations are quite demanding in computational time and resources the vorticity part is considered as the only source of fluid stress for the purpose of computation cost reduction. In fact, fluid shear stress is contributed by both strain and vorticity. In mathematics, the computation of stress can be performed by strain only, vorticity only, or both. The computational results are exactly the same. The NSE equation adopts strain, which is symmetric and stress based on Stokes's assumption. In [1], a new governing equation which is based on a new assumption that accepts that fluid stress has a linear relation with vorticity, which is anti-symmetric. According to the mathematical

analysis, the new governing equation is identical to NS equations in numerical analysis, but in a physical sense, the new governing equation is just the opposite to NSEs as it assumes that fluid stress is proportional to vorticity, where both are anti-symmetric, but not strain, contrary to Stokes's assumption and the current NSE.

Although both NSEs and the new governing equation in [1] lead to the same computational results for laminar flow, the new governing equation has several advantages: 1. The vorticity tensor is anti-symmetric, which has three elements, but NSEs use the strain tensor, which has six elements. It is shown that the computational cost is reduced to half for the viscous term. 2. The anti-symmetric matrix is independent of the coordinate system change or Galilean invariant, but the symmetric matrix that NSE uses is not. 3. The physical meaning is clear that the viscous term is generated by vorticity, not by strain only. 4. The viscosity is obtained by experiments, which are based on vorticity but not strain, since both strain and stress are hard to measure experimentally. 5. Vorticity can be further decomposed to rigid rotation and pure anti-symmetric shear, which is very useful for further study turbulent flow. However, the NS equation has no vorticity term, which is an impediment for further turbulence research. ??ref [27] in [1] studied the mechanism of turbulence generation and concluded that shear instability and transformation from shear to rotation are the paths of flow transition from laminar flow to turbulent flow. Using Liutex and the third generation of vortex identification methods, a lot of new physics has been found (see Dong et al., ??iu et [1]) In Ref.28 in [1], Zhou et al. elaborated the hydrodynamic instability induced turbulent mixing in wide areas, including inertial confinement fusion, supernovae, and their transition criteria. Since the new governing equation has a vorticity term, which can be further decomposed to shear and rigid rotation, the new governing equation would be helpful in studying flow instability and transition to turbulence. Turbulence is rotational and characterized by large fluctuations in vorticity and thus it is important to accurately define vorticity. In the vorticity equation the vortex stretching term can be argued to be one of the most important mechanisms in the turbulence dynamics. It represents the enhancement of vorticity by stretching and is the mechanism by which the turbulent energy is transferred to smaller scales.

The purpose of this article is to refer to the periodic NS equations with high energy assumption as in the case of the continuum hypothesis being valid and can breakdown in finite time but with sufficient low energy scaling as in a fractal setting like for example on a Cantor set, the equations may not exhibit finite time blowup. It is known recently in the literature that the Cantor set with layers N (N can have up to two orders of magnitude) can be presented as a potential contender (analytical framework) for connecting the energy in a molecular level say 1 at some cutoff length scale l to the energy at a continuum level L with length scale L . The equipartition theorem of statistical mechanics has been used (Terrence Tao 2015) to relate the energy of a discrete block in say 1 (molecular scale) to the energy in L (continuum scale). Additionally it has been shown that the ratio of the energy of the continuum scale to the molecular scale is a factor of 2^N . It then makes intuitive sense that the high energy PNS problem may breakdown in finite time. This article gives a general model using specific periodic special functions, that is degenerate elliptic Weierstrass P functions. See Figure 1.

The definition of vorticity should be as defined in [1], which is that vorticity is a rotational part added to the sum of antisymmetric shear and compression and stretching. A vortex is recognized as the rotational motion of fluids. Within the last several decades, a lot of vortex identification methods have been developed to track the vortical structure in a fluid flow; however, we still lack unambiguous and universally accepted vortex identification criteria. It has been uncovered that the regions of strong vorticity and actual vortices are weakly related. It recently [1] has been concluded that a vorticity vector does not only represent rotation but also claims shearing and stretching components to be a part of the vortical structure, which is contaminated by shears in fluid. Satisfying a divergence free vector field and periodic boundary conditions respectively with a general spatio-temporal forcing term $\partial_t \mathbf{u} = \mathbf{f}(\mathbf{x}, t)$ which is smooth and spatially periodic, the existence of solutions of PNS which blowup in finite time can occur starting with the first derivative and higher with respect to time. On the other hand if $\mathbf{f}$ is not smooth, then there exist globally in time solutions on $[0, \infty)$ with a possible blowup at $t = \infty$. The control of turbulence is

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3 II. Materials and Methods

Consider the incompressible 3D Navier Stokes equations defined on the three-Torus $\mathbb{T}^3 = \mathbb{R}^3 / \mathbb{Z}^3$. The periodic Navier Stokes system is,
$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} - \nabla \cdot \mathbf{\tau} = \mathbf{f} + \nabla \pi, \quad \text{div } \mathbf{u} = 0$$
 where $\mathbf{u} = (u, v, w)$ is velocity, π is pressure and $\mathbf{f} = (f, g, h)$ is forcing vector. Here $\mathbf{\tau} = \mu \nabla^2 \mathbf{u} + \mu \nabla \mathbf{u} \cdot \nabla$, where μ is viscosity, and $\nabla \mathbf{u} \cdot \nabla$ denote respectively the ∂_x, ∂_y and ∂_z components of velocity.

Introducing Poisson's Equation (see [2], [3] and [5]), the second derivative $\nabla^2 \mathbf{u}$ is set equal to the second derivative obtained in the $\nabla \mathbf{u} \cdot \nabla$ expression further below, as part of $\mathbf{f}$, and $\nabla^2 \mathbf{u} = \nabla \mathbf{u} \cdot \nabla$, $\nabla^2 \pi = \nabla \mathbf{u} \cdot \nabla \cdot \mathbf{u}$. Furthermore the right hand side of the one parameter group of transformations are next mapped to $\mathbf{u}$ variable terms, (note that $\mathbf{u}$ and π are not assumed to be arbitrarily small, they can be at

most order one), $?? \text{ } ??? = -?? \text{ } ?? \text{ } ?? \text{ } ? * = 1 \text{ } ?? \text{ } ?? \text{ } ?? \text{ } , ?? * = 1 \text{ } ?? \text{ } 2 \text{ } ?? , ?? \text{ } ?? * = ??? \text{ } ?? \text{ } , ?? * = ?? \text{ } 2 \text{ } ?? , ?? = 1,2,3.$

[illegible]

possible to maintain when the initial conditions and boundary conditions are posed properly for (PNS) ([5]). The endpoint regularity in Onsager's conjecture is addressed, and it is found that conservation of energy occurs when the Hölder regularity is exactly 1/3. Finally it is proposed that the periodic Liutex new equations [1] (The new equations referred to previously) do not exhibit finite time blow up. This is the focus of the ongoing work of the author to be presented in the near future.

Year 2023 () I??(??) ??1 = 1 ?? 6 ? ? ? ? ? ? ? ? (?? -1 -1) ? ??? 3 ??? ? 2 + ?? ? ??? 3 ??? ? ? ? 2 ?? 3 ??? 1 2 + ? 2 ?? 3 ??? 2 2 + ? 2 ?? 3 ??? 3 2 ? ?? + (?? -1 -1) ? ??? 3 ??? ? ??? ??? 3 ?? ? ? ? ? ? ? (1 -? -1) ??(??) ??4 = 1 ?? 3 ??? 2 ?? ?? ????? ? ? ?? 1 ?? 2 ?? 3 2 -?? 3 ?? 3 ??? 3 ??? 3 ?? ?? + ?? 2 ?? ? ? ?(?? 3 ?? ??)?

It has been shown in Moschandreou et al [5] that this decomposition holds and that, $\chi_1 + \chi_2 + \chi_3 + \chi_4 = 3 \chi_5$

The function $\mathcal{F}(\mathbf{r})$ is the surface integral of pressure terms minus the volume integral of tensor product term.

At the end of this paper, a proof that on a volume of an arbitrarily small sphere embedded in each cell of the lattice centered at $(\frac{1}{2}, \frac{1}{2}, \frac{1}{2})$ (centers of cells) we have, $\langle \sigma_i \sigma_j \rangle = \langle \sigma_i \sigma_k \rangle = \langle \sigma_i \sigma_l \rangle = 0$

From this equation we then can solve for ψ algebraically and differentiating with respect to x and using Poisson's equation (setting the representation of each of the two partial derivatives with respect to x equal to each other we obtain, $\psi = 0$, which is precisely the following PDE, $\nabla^2 \psi = -\rho$)

$$\begin{aligned} & \frac{\partial}{\partial x} \left(\epsilon_0 \frac{\partial \psi}{\partial x} \right) + \frac{\partial}{\partial y} \left(\epsilon_0 \frac{\partial \psi}{\partial y} \right) = -\rho \\ & \epsilon_0 \frac{\partial^2 \psi}{\partial x^2} + \epsilon_0 \frac{\partial^2 \psi}{\partial y^2} = -\rho \end{aligned}$$

where ρ is the charge density.

where $\delta \text{ } ?? \delta \text{ } ?? \text{ } ? = ??? \text{ } ?? \text{ } 1 , \text{ } ?? \text{ } ?? \text{ } 2 , \text{ } ?? \text{ } ??? \text{ } ?$ is the forcing vector and $?? \text{ } ? = (?? \text{ } 1 , ?? \text{ } 2 , ?? \text{ } 3)$ is the velocity in each cell of the 3-Torus.

For the three forcing terms, set them equal to products of reciprocals of degenerate Weierstrass P functions shifted in spatial coordinates from the center (?? ?? , ?? ?? , ?? ??), ?? = 1 . ??.

Here the $(\frac{1}{2}, \frac{1}{2}, \frac{1}{2})$ is the center of each cell of the lattice belonging to the flat torus. Upon substituting the Weierstrass P functions and their reciprocals (unity divided by P-function) into Eq.(1) together with the forcing terms given by $\vec{F}$, it can be observed that in the equation that terms in it are multiplied by reciprocal Weierstrass P functions which touch the centers of the cells of the lattice, thus simplifying Eq.(1). The initial condition in $\vec{F} = 0$ is instead of a product of reciprocal degenerate Weierstrass P functions for forcing, is a sum of these functions. The parameter $\vec{F}$ in the degenerate Weierstrass P function, if chosen to be small gives a ball, $\vec{F} = \{F_1, F_2, F_3 : F_1^2 + F_2^2 + F_3^2 = (F_1^2 + F_2^2 + F_3^2)^{1/2}\}$

Here we are in Cartesian space $\nabla \cdot \mathbf{u} = 0$ with 2-norm $\|\mathbf{u}\|_2$. Since the terms are squared in length in the initial condition for $\nabla \cdot \mathbf{u} = 0$ we require to multiply by dynamic viscosity μ to obtain units of velocity. In the above, the forcing is taken to be different than the gradient of pressure.

Introducing the space $\varphi(??\;3,\; ??) = ???\; ?\; ? + , \; ??\; 3\; ?\; ??\; ???\; 3\; ??\; ??\; ;\; ???\; : \; 2??\; 1\; ??\; 1 + ??\; 2 = 0 \&????$
 $1 + ?????\; 2 + ?? = 0,$ $??? \; 1, \; ??\; 2\; ?\; ?? \times ??\; (??\; ?\; ?)\&??\; 2 = ??\; 1\; 2 \&\; ??\; 3\; (??\; 1, \; ??\; 2, \; ??\; 3, \; ??)\; ?\; ??\; 0$
 $(??\; 3)\;),$

where $?? \text{ } ?? \text{ } ?? = (?? - 1)?? \text{ } 1 \text{ } ???? \text{ } 3 \text{ } ???? + 2???? \text{ } 3 \text{ } ???? \text{ } 1 \text{ } ???? \text{ } ? \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 1 + (?? - 1)?? \text{ } 2 \text{ } ???? \text{ } 3 \text{ } ???? + 2???? \text{ } 3 \text{ } ???? \text{ } 2 \text{ } ???? \text{ } ? \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 2 - ???? \text{ } 3 \text{ } ???? \text{ } ??? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 1 \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 1 \text{ } ???? \text{ } 3 \text{ } ???? + ?? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 2 \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 2 \text{ } ???? \text{ } 3 \text{ } ???? - ???? \text{ } 3 \text{ } ???? \text{ } 1 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } - ???? \text{ } 3 \text{ } ???? \text{ } 2 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 2 \text{ } ???? \text{ } ? + ?? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 1 \text{ } ???? \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 3 \text{ } ???? ???? \text{ } 3 \text{ } 2 \text{ } ?? = (?? - 1) \text{ } ?? \text{ } 1 \text{ } ???? \text{ } 3 \text{ } ???? + 2???? \text{ } 3 \text{ } ???? \text{ } 1 \text{ } ???? \text{ } ? \text{ } ?? \text{ } 2 \text{ } ?? \text{ } 3 \text{ } ???? \text{ } 3 \text{ } ???? \text{ } 1 + (?? - 1) \text{ } ?? \text{ } 2 \text{ } ???? \text{ } 3 \text{ } ???? + 2???? \text{ } 3 \text{ } ???? \text{ } 2$

$\begin{aligned}
& \text{???? ? ?? 2 ?? 3 ???? 3 ???? 2 + ???? 3 ???? ? ???? 3 ???? 1 ???? 3 ???? 3 ???? 1 ???? + ???? 3 ???? 2 ????} \\
& \text{3 ???? 3 ???? 2 ???? ? + ? ???? 2 ???? ? 2 ?? 3 ???? 3 ???? 2 ? ???? 3 ???? ???? 2 ???? ? ?? 3 +? ???? 1} \\
& \text{???? ? 2 ?? 3 ???? 3 ???? 1 ? ???? 3 ???? ???? 1 ???? ? ?? 3}
\end{aligned}$

Here it is clear that there exists a solution of PNS that is not smooth in time ?? for the first and higher derivatives.

6 The Full Equation Proof for the Periodic Navier Stokes Equations

Integrating the Navier Stokes equations over an epsilon ball we obtain, $\int_{B_\epsilon} \rho \frac{Dv}{Dt} = -\int_{B_\epsilon} \nabla p + \int_{B_\epsilon} \nabla \cdot \tau + \int_{B_\epsilon} \rho f$ (IV)

The first part of ?? 3 becomes, ?? 3 = ? ? ?(?),

where $\varphi_i = 1$ if $i \in \{1, 2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12, 13, 14, 15, 16, 17, 18, 19, 20, 21, 22, 23, 24, 25, 26, 27, 28, 29, 30, 31, 32, 33, 34, 35, 36, 37, 38, 39, 40, 41, 42, 43, 44, 45, 46, 47, 48, 49, 50, 51, 52, 53, 54, 55, 56, 57, 58, 59, 60, 61, 62, 63, 64, 65, 66, 67, 68, 69, 70, 71, 72, 73, 74, 75, 76, 77, 78, 79, 80, 81, 82, 83, 84, 85, 86, 87, 88, 89, 90, 91, 92, 93, 94, 95, 96, 97, 98, 99, 100\}$ and $\varphi_i = 0$ otherwise.

where p is the pressure and ρ is the density of the fluid.

Dividing Eq.(IV) by the measure or volume of the ball of radius epsilon centered at point ??.

we know since ϕ is continuous everywhere on the 3-Torus (since integrals are continuous in inverting gradient), and in particular at the the center of the epsilon ball (note higher order derivatives of ϕ 3 blowup, not ϕ 3 and pressure), then, (V) However using the Fet theory, we can see that the integral on the RHS of Eq.(IV) divided by the volume of the ball is related to the integral over the ball centered at $\mathbf{x}_i$ of Eq.(VI) is the PDE we obtained previously and occurs at an arbitrarily small epsilon ball centered at each cell of the lattice of the 3-Torus.

In reference [5], we showed that, where? $1 = ? + ? 2$

Recall that the three velocities are isotropic and they are continuous on ∂B_ϵ and ψ is continuous on the epsilon ball. Also ψ is independent of r for Hölder continuous functions at $r = 1/3$. Also if we specify the time, the solution is a Hölder continuous function in the data ψ with a Hölder constant equal to one half. Since the negative pressure gradients are greater than or equal to zero being reciprocal Weierstrass P functions and $\psi_3 \geq 0$ and ψ_1 and ψ_2 cancel in the space (ψ_3, ψ_1) when integrating on the six faces of surface of a cell of $\mathbb{R}^3$, we have that, Theorem 1.1 + 1.2 + 1.4 = 0 if and only if ψ is continuous on the epsilon ball ∂B_ϵ .

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Proof: Apply (V) to ? 1

8 IV. Conclusion

Satisfying a divergence free vector field and periodic boundary conditions respectively with a general spatiotemporal forcing term $\delta \varphi$ which is smooth and spatially periodic, the existence of solutions which blowup in finite time for PNS can occur starting with the first derivative and higher with respect to time. P. Isett (2016) (see [13]) has shown that the conservation of energy fails for the 3D incompressible Euler flows with Hölder regularity below $1/3$. (Onsager's second conjecture) The endpoint regularity in Onsager's conjecture has been addressed, and it is found that conservation of energy occurs when the Hölder regularity is exactly $1/3$. The solution for Euler's equation given in this paper agrees with this fact that gradient of φ with respect to spatial position φ does in fact diverge. This is a short-distance/ultraviolet (UV) divergence in the language of quantum field-theory as L. Onsager proposed. Finally very recent developed new governing equations of fluid mechanics are proposed to have no finite time singularities. This is the focus of the ongoing work of the author to be presented in the near future. Finally future work to conclude the nature of flows in a non-epsilon or arbitrary small ball for the 3-Torus will be carried out. ^{1 2}

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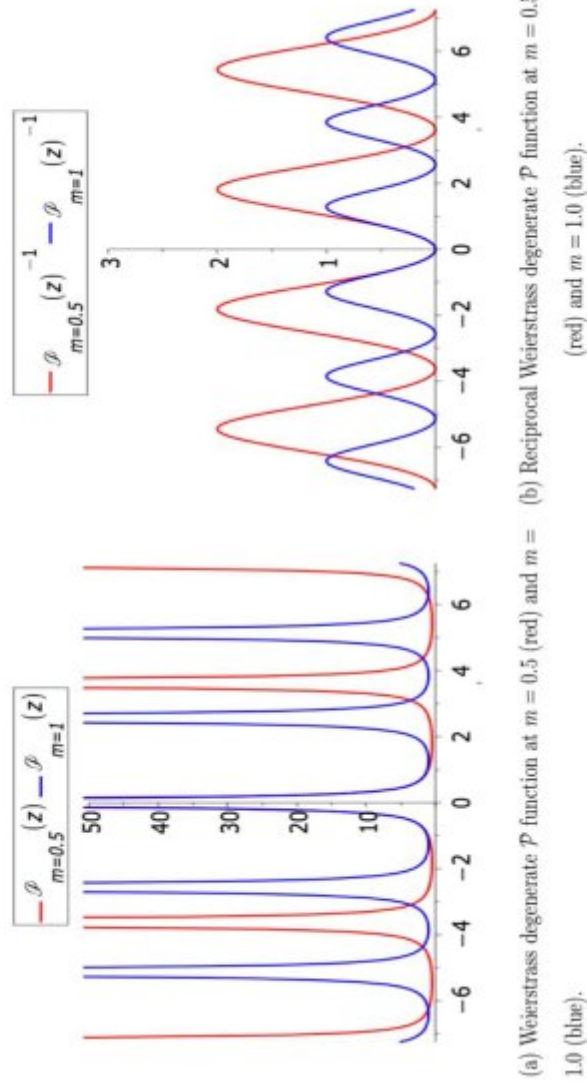
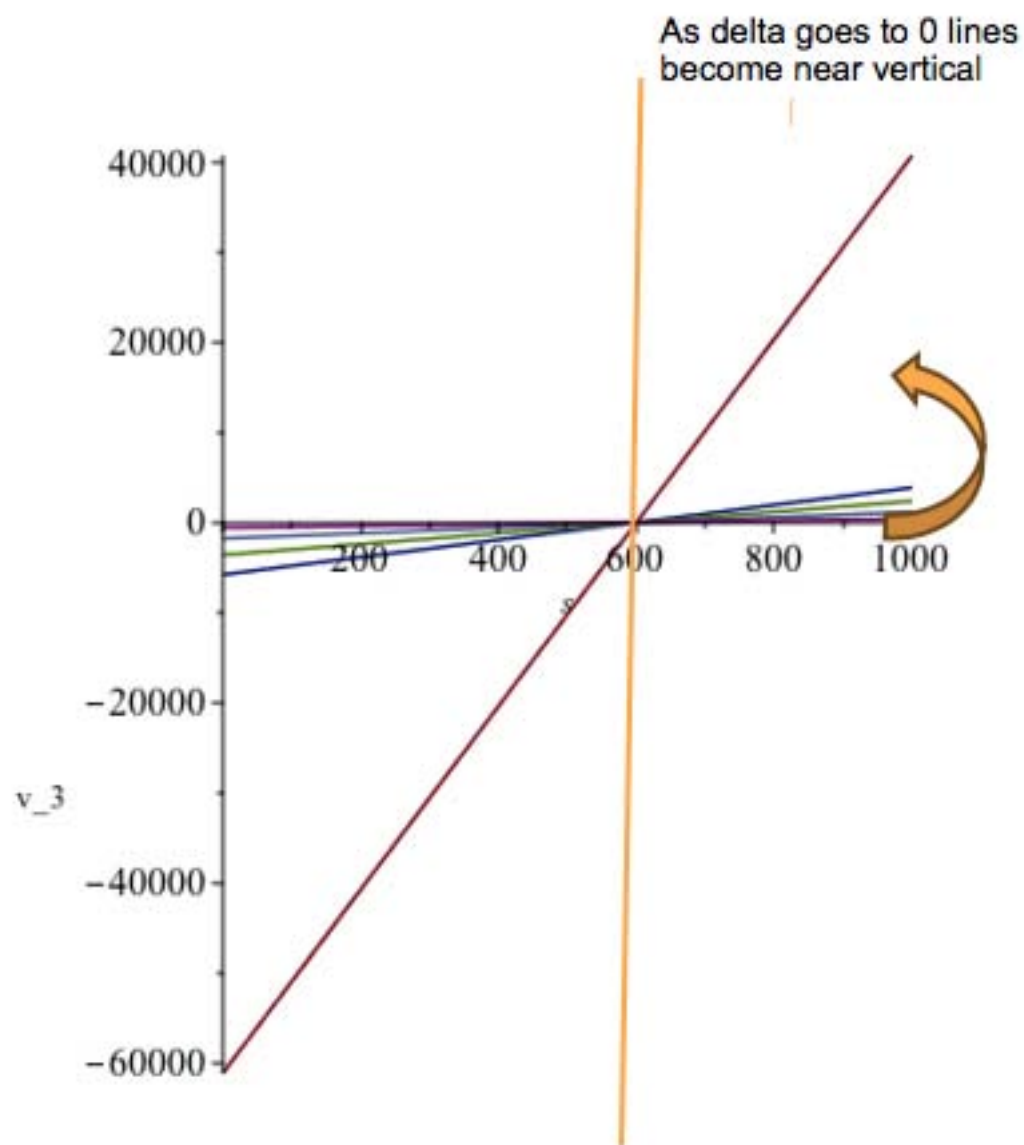


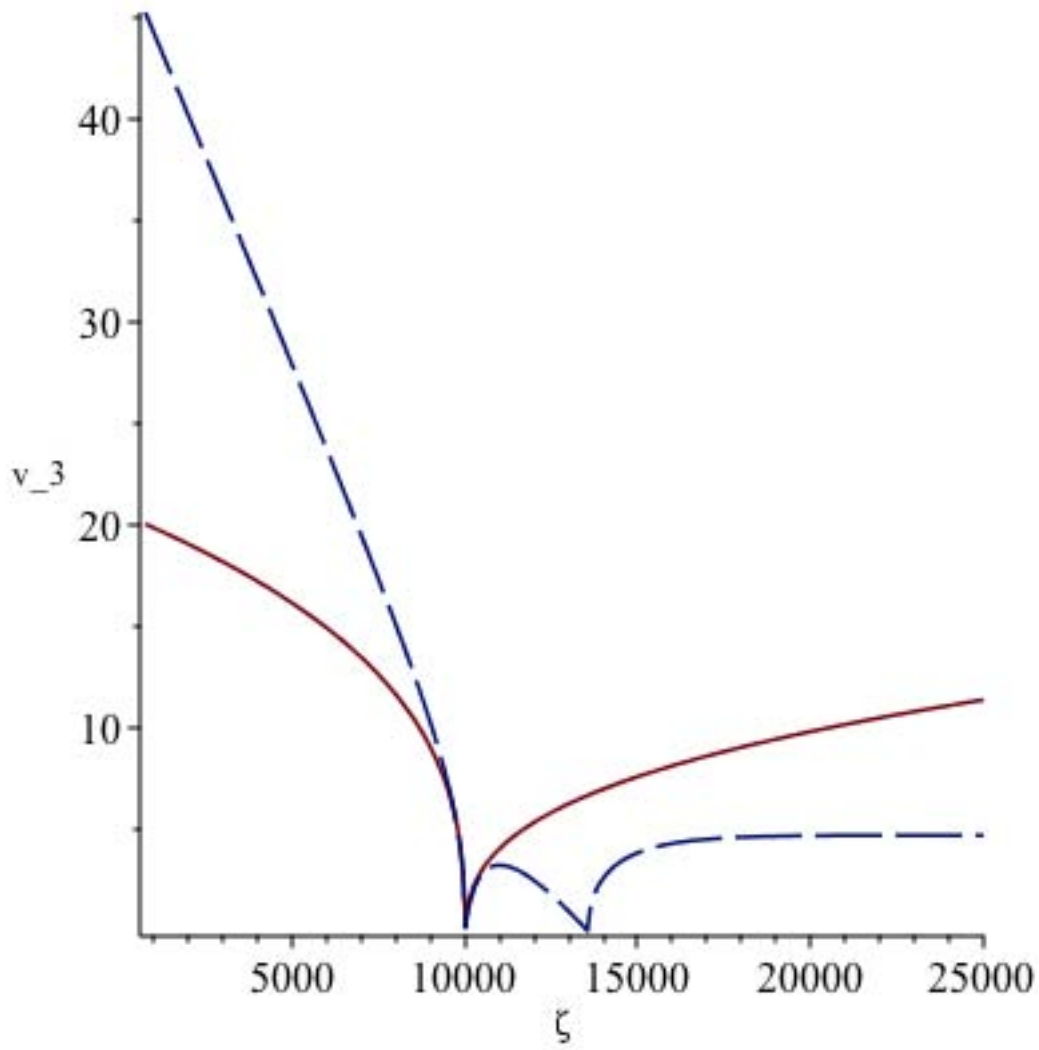
Figure 1: Plots of the Degenerate Weierstrass P functions, given relative to the canonical Weierstrass P functions $\wp(z, g_2, g_3)$ as $\mathcal{P}_{m=n}(z) = \wp(z, 3n^2, n^3)$. Here, displayed as the standard function (Fig. 1a), and its reciprocal (Fig. 1b).

Figure 1: Figure 1c :



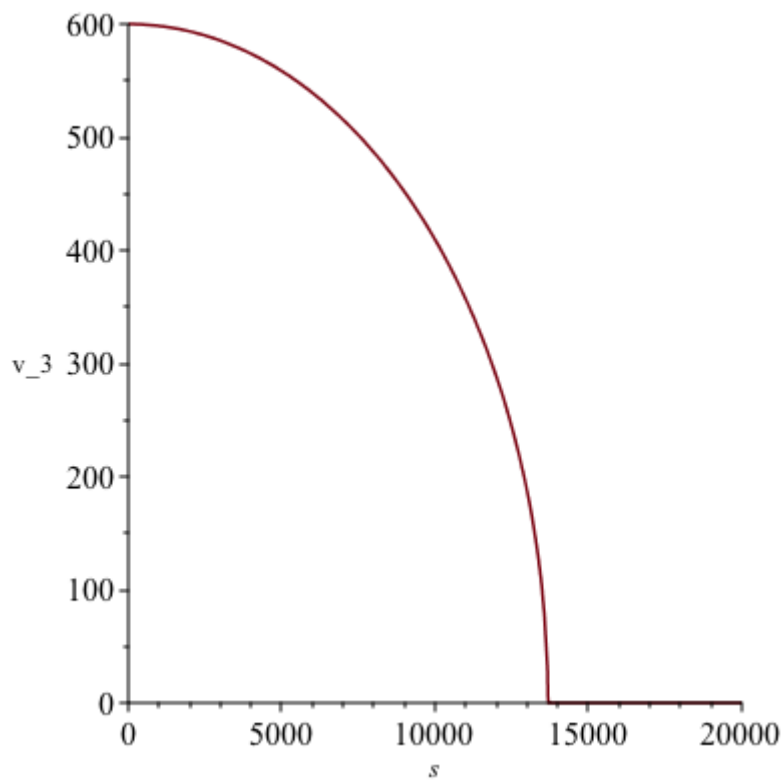
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Figure 2: Figure 2 :



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Figure 3: ???? 3 ????



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Figure 4: Figure 4 :

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Figure 5:

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Figure 7: I

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Figure 8:

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